



Hubble High-contrast image processing & Debris Disk imaging

Elodie Choquet

1. Make Hubble as competitive as SPHERE

SPHERE: The State-of-the-art in HC imaging

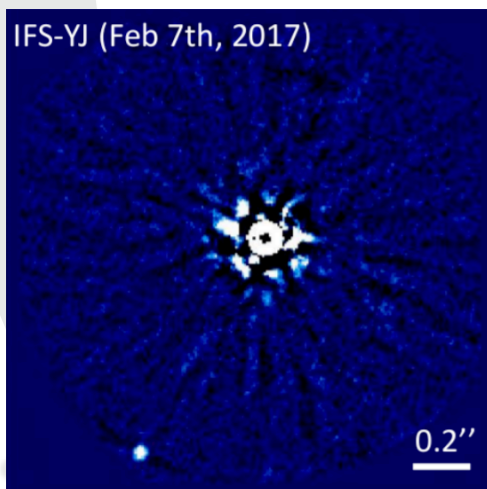
Highly Optimized for planets hunt:

- ▶ Extreme AO system
- ▶ Stabilized platform
- ▶ Optimized coronagraphs

Deepest observations ever achieved:

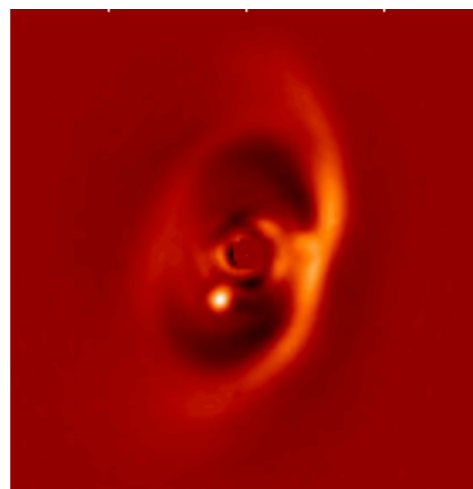
Beuzit et al. 2019

6-12 M_{jup} planet

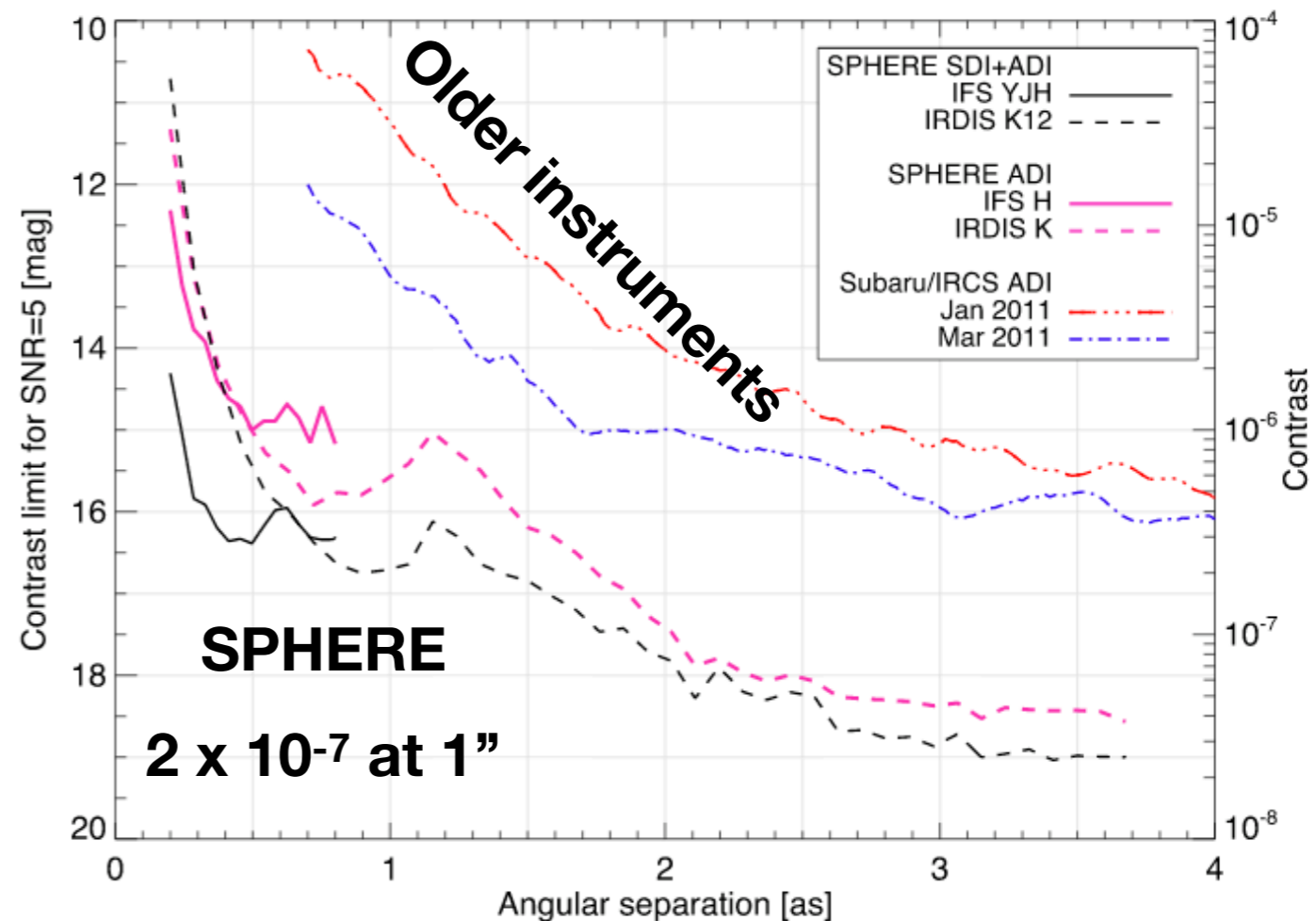


Chauvin et al. 2017

2-17 M_{jup} planet



Müller et al. 2018



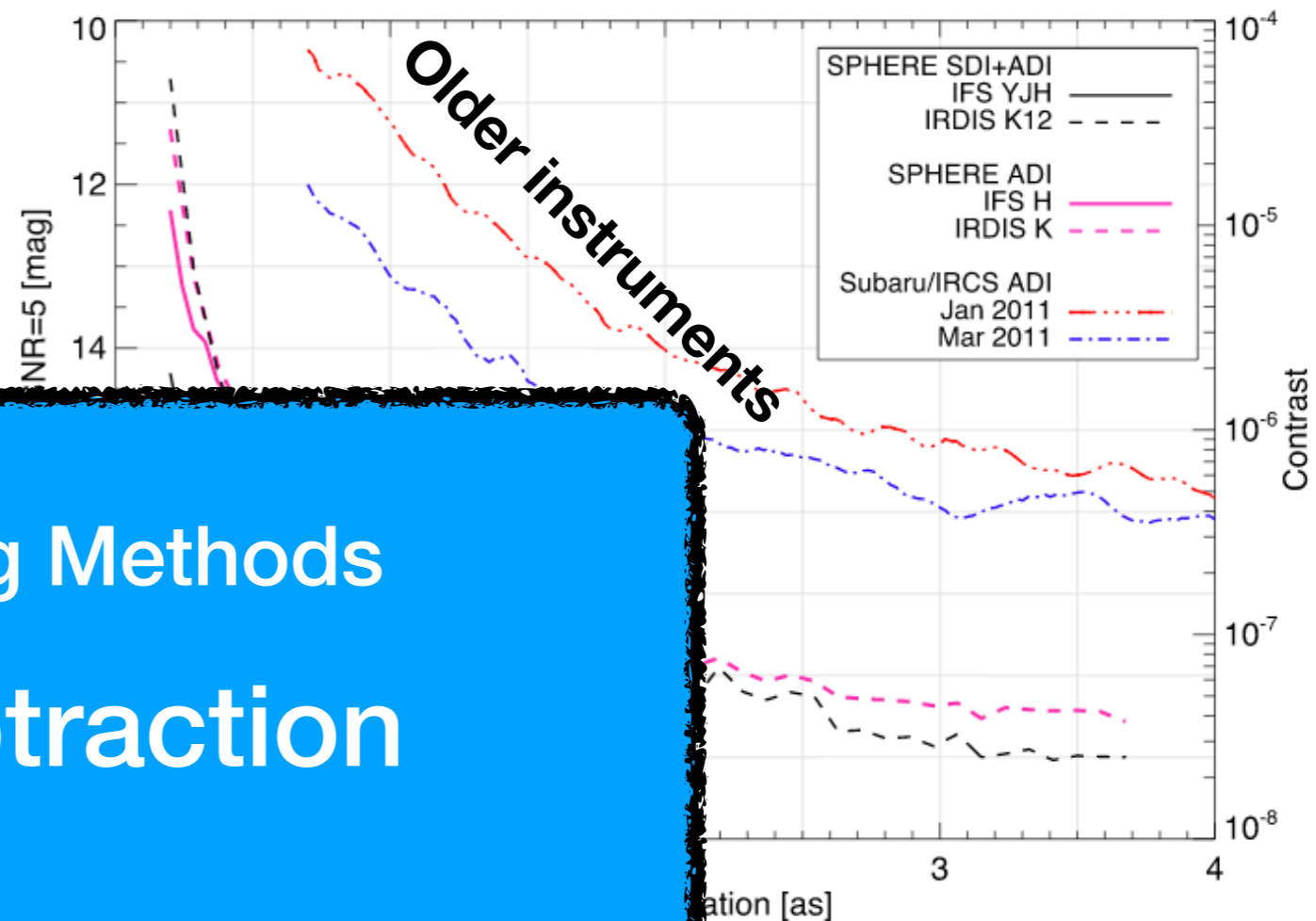
Vigan et al. 2015

SPHERE: The State-of-the-art in HC imaging

Highly Optimized for planets hunt:

- ▶ Extreme AO system
- ▶ Low noise detectors
- ▶ Optimized coronagraphs
- ▶ Stabilized platform

Deepest observations ever achieved:



▶ Optimized data processing Methods

Stellar PSF Subtraction



Chauvin et al. 2017



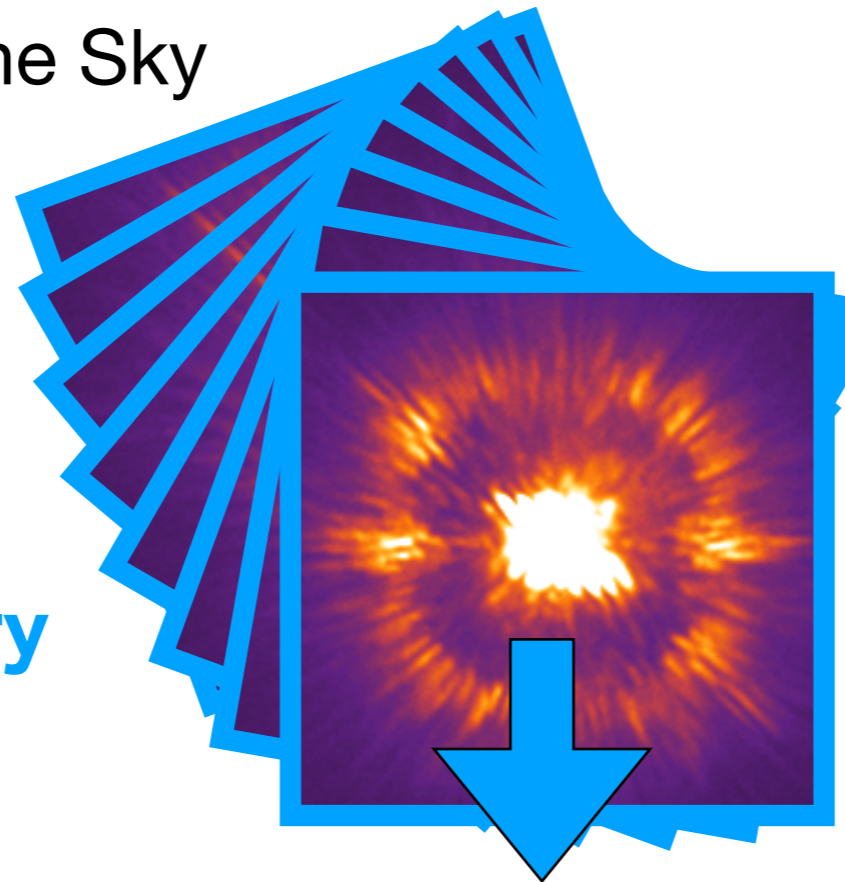
Müller et al. 2018

Vigan et al. 2015

Well-established method

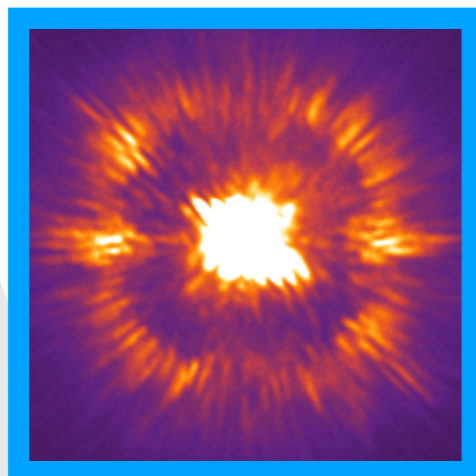
Natural Rotation of the Sky

PSF Library

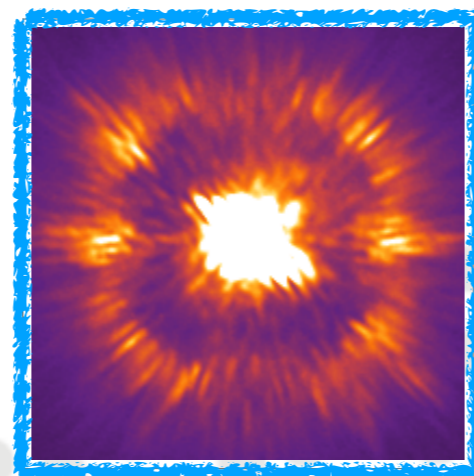


Linear combination
Median, ...

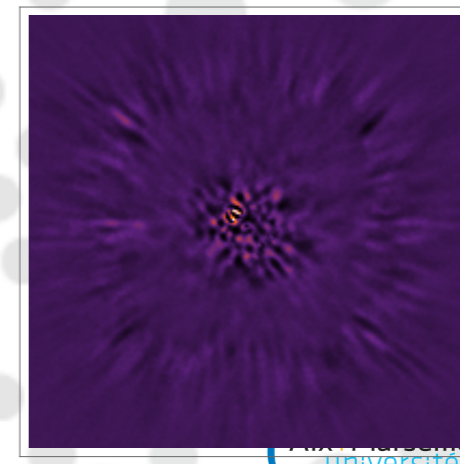
Science Image



Model



Residuals

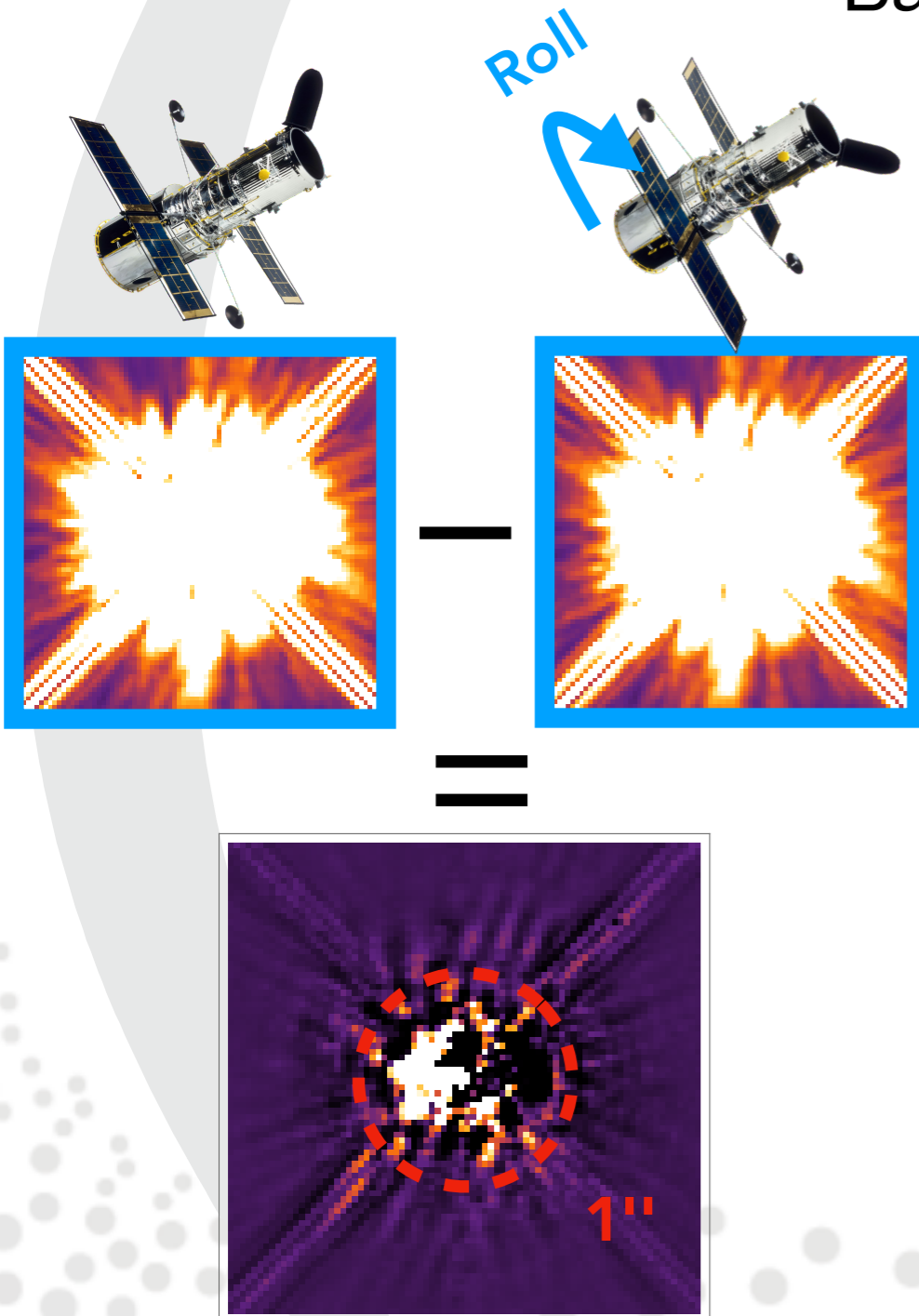


-

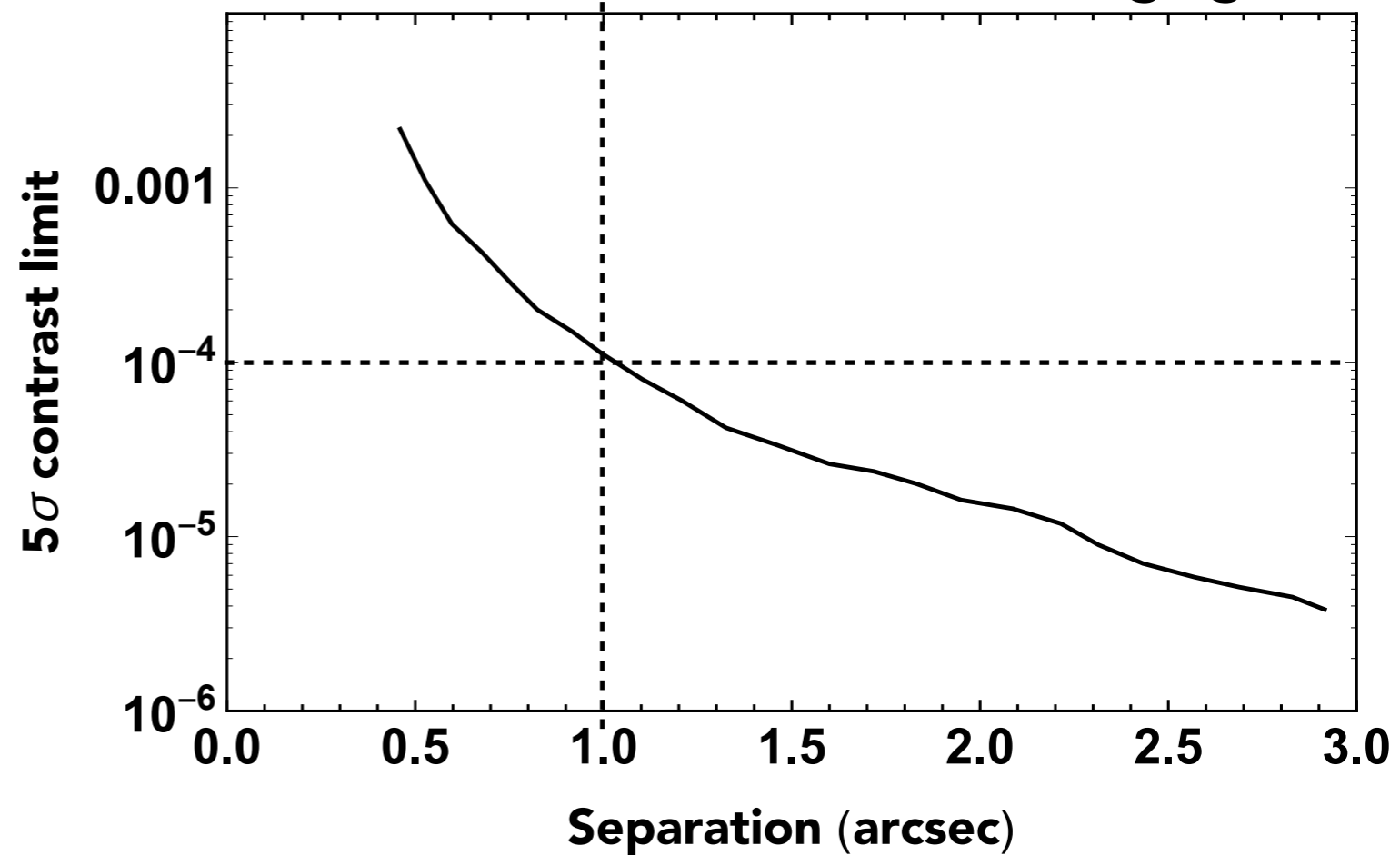
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High-contrast Imaging with Hubble

Basic differential imaging



Median Limits Differential Imaging

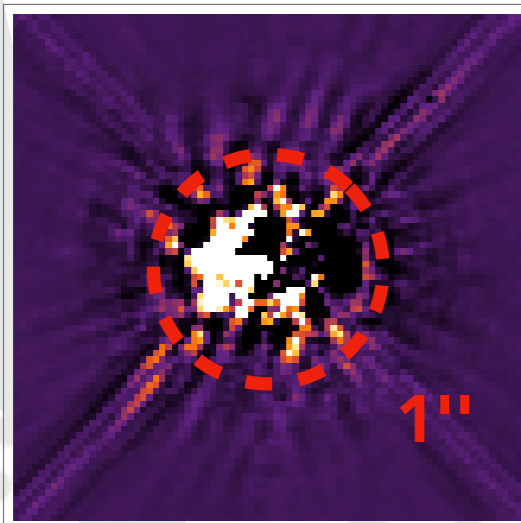
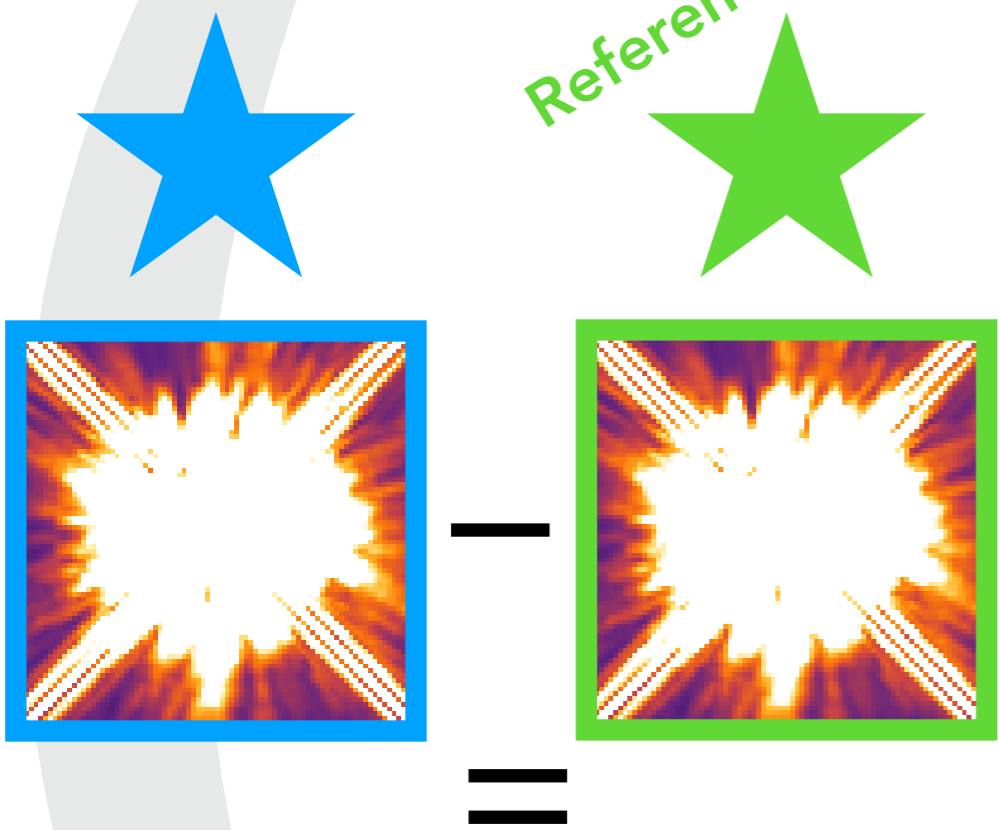


Lowrance et al. 2005

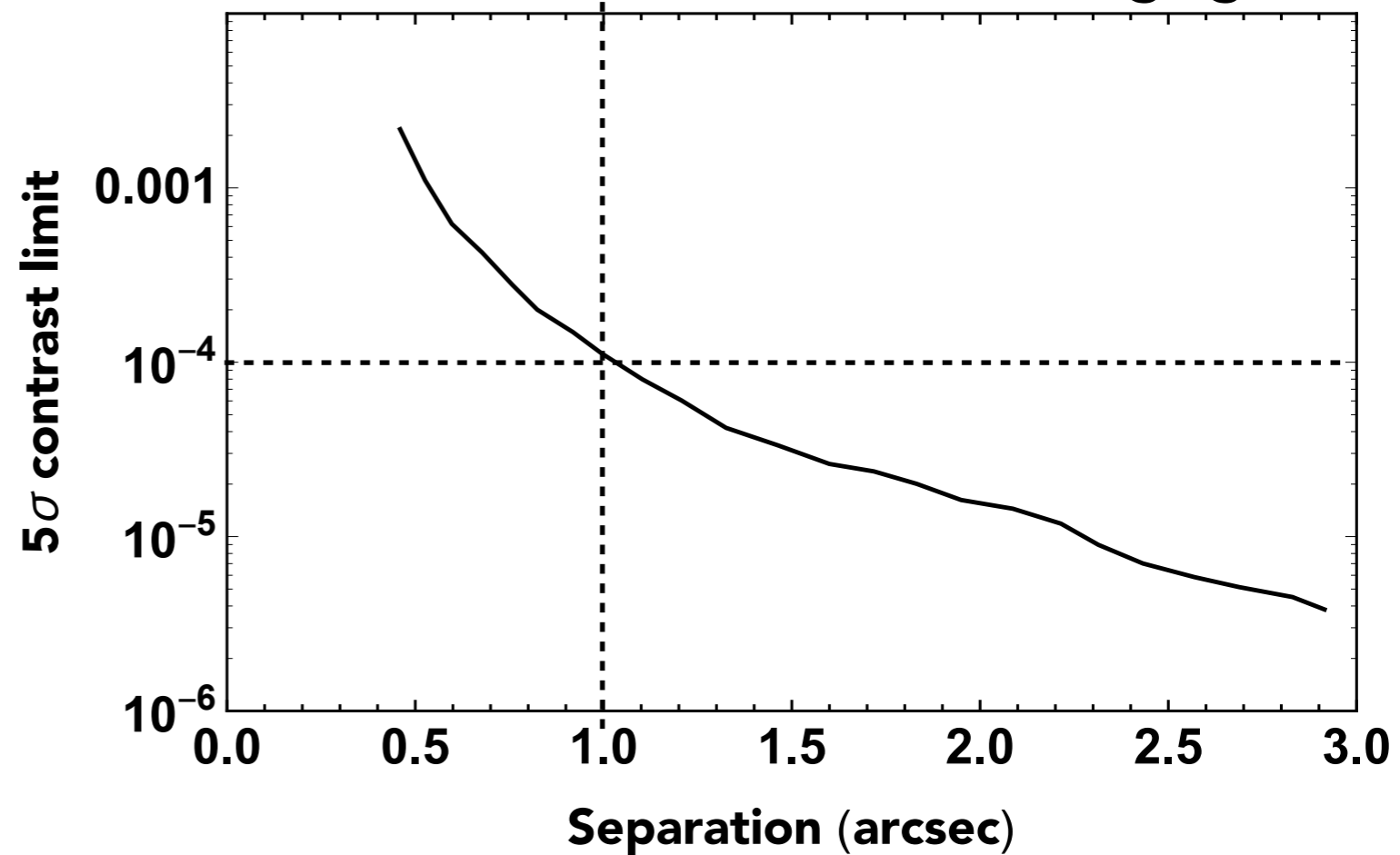
High-contrast Imaging with Hubble

Basic differential imaging

Reference



Median Limits Differential Imaging



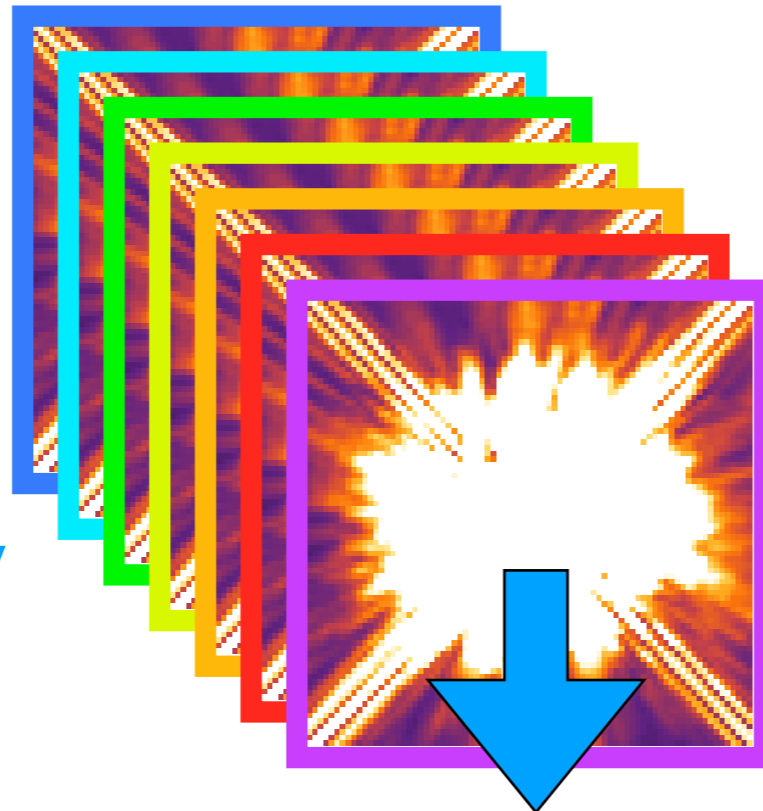
Lowrance et al. 2005

A Novel Method

Multiple Reference Stars

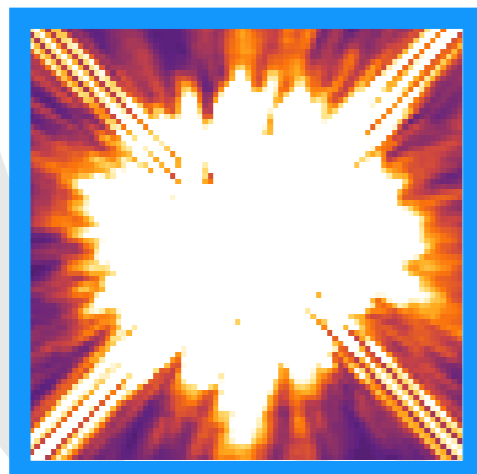
- ▶ 10^2 Reference stars
- ▶ 10^3 PSF images
- ▶ Samples all variations

Archival Library

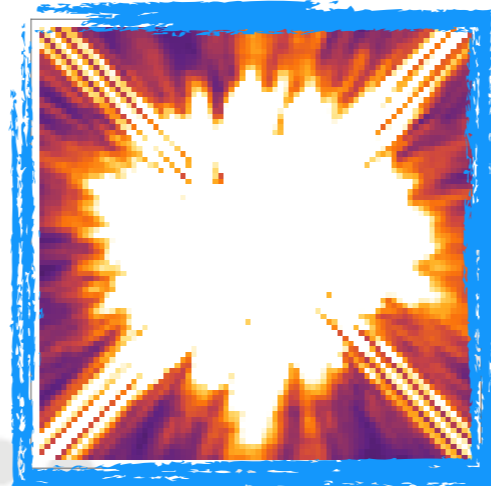


Modern Algorithms

Science Image



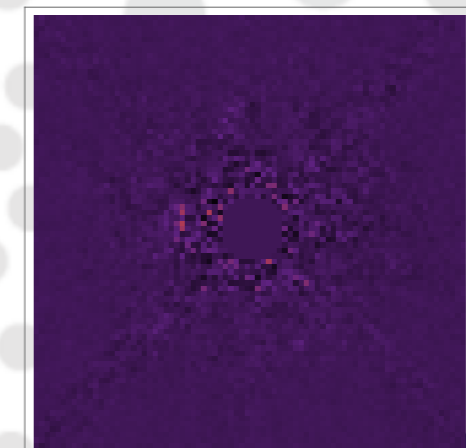
Model



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Residuals



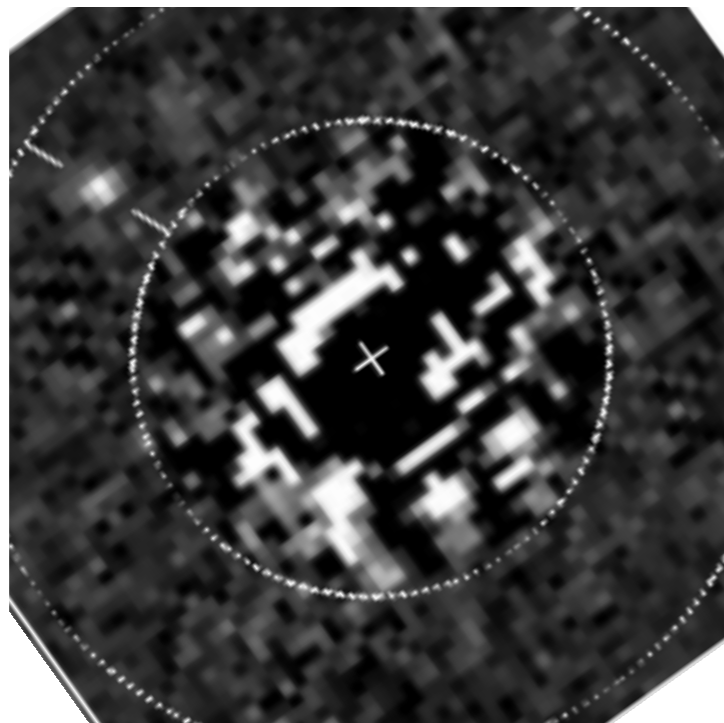
First demonstration: HR 8799 system

Archival library

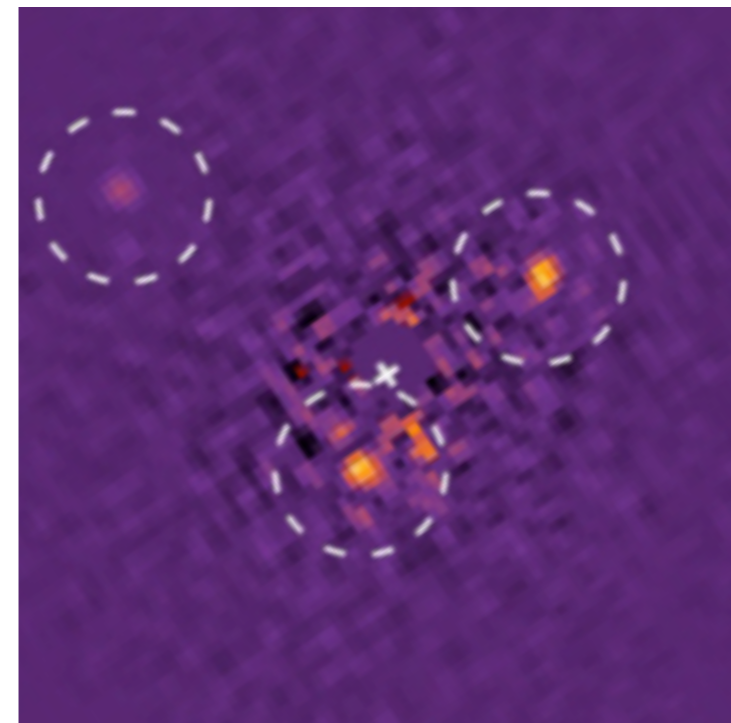
- ▶ 23 Reference stars
- ▶ 200 PSF images

1998 HST data

Lafrenière et al. 2009



Soummer et al. 2011

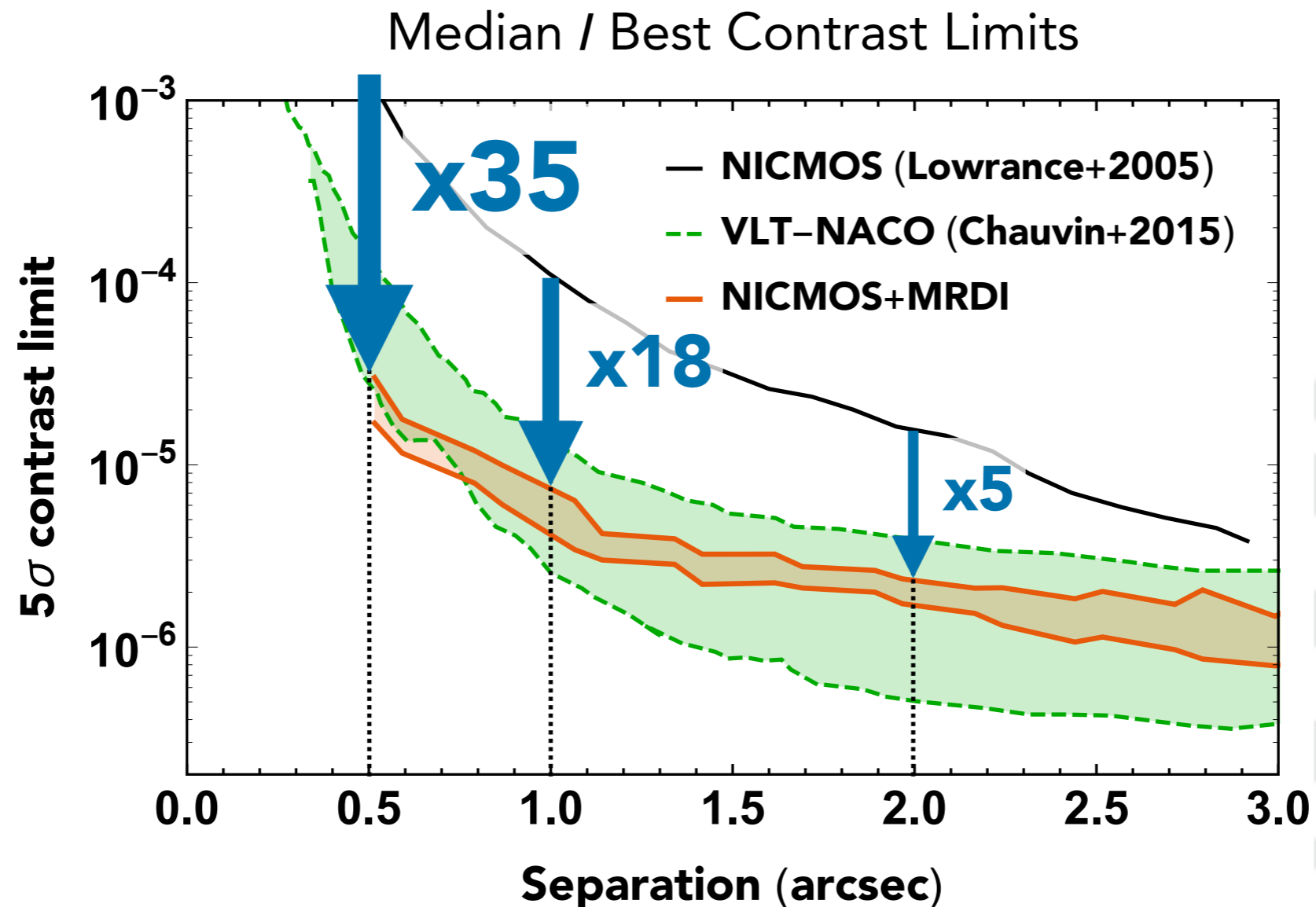


ALICE: Optimization & Generalization

Archival library

- ▶ 80 Reference stars
- ▶ 850 PSF images

Full NICMOS archive



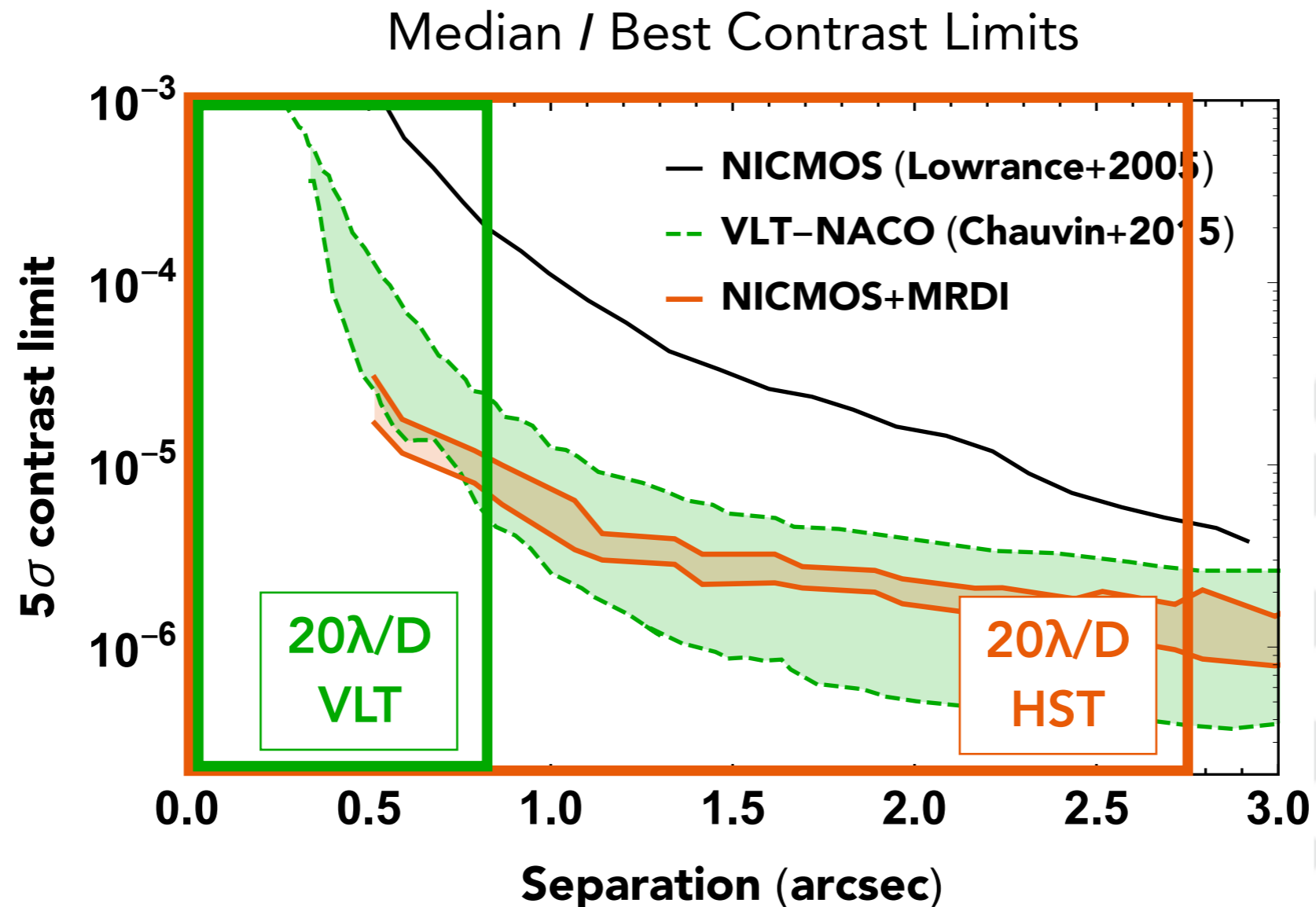
Soummer et al. 2012
Choquet et al. 2014
Hagan et al. 2018

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Full NICMOS archive



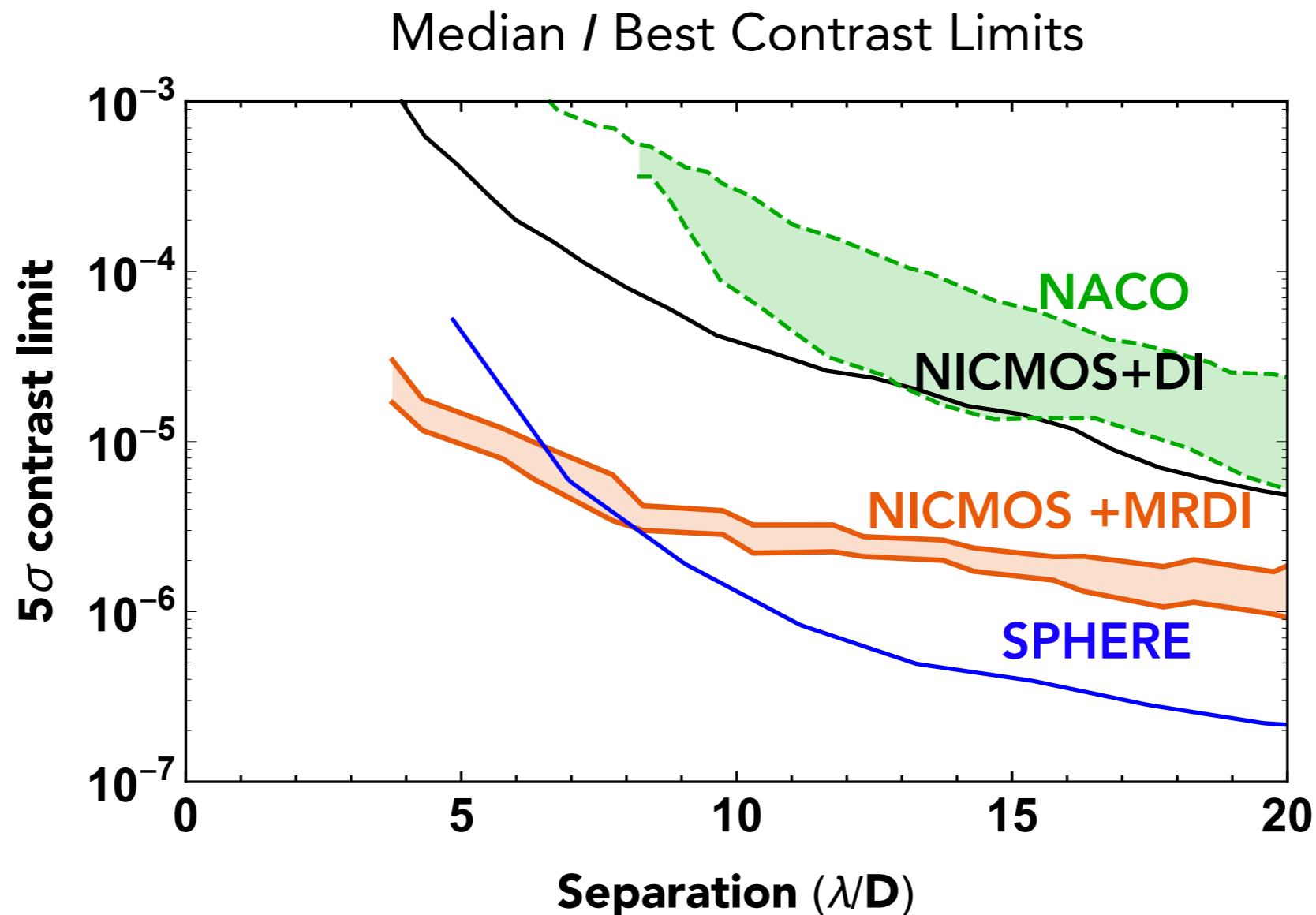
Soummer et al. 2012
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ALICE: Optimization & Generalization

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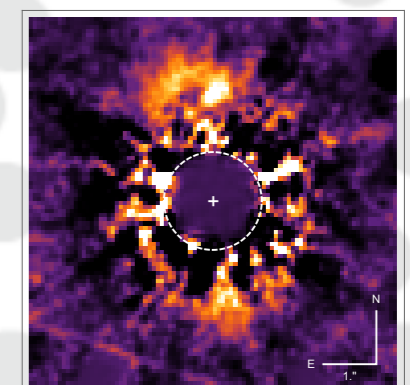
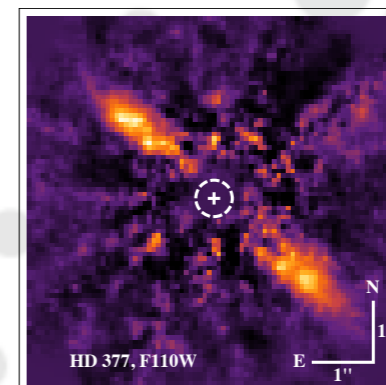
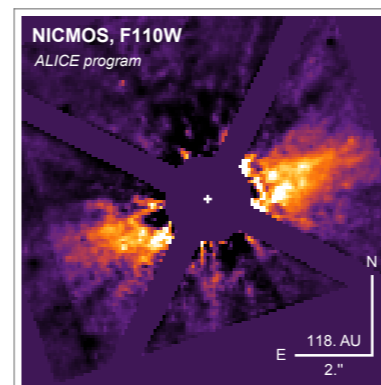
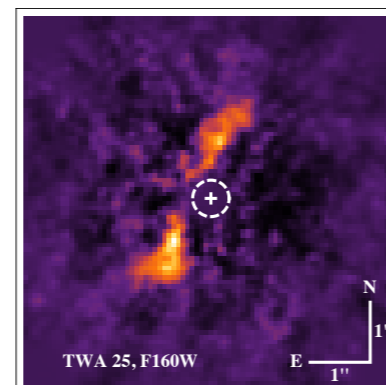
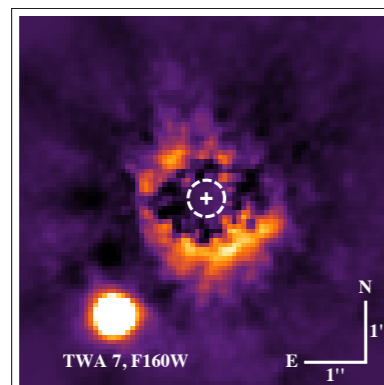
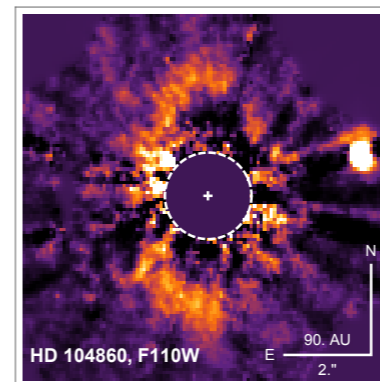
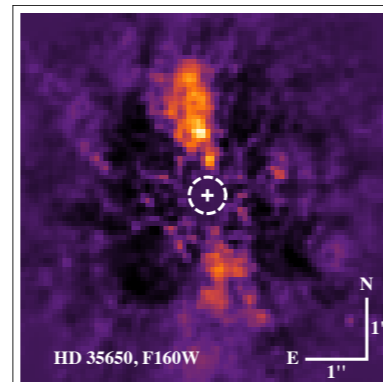
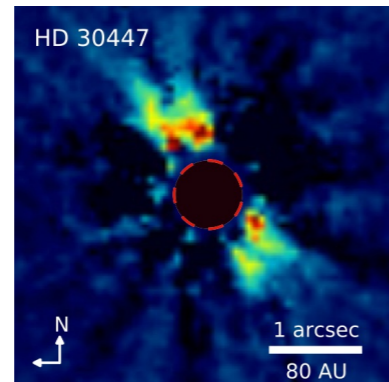
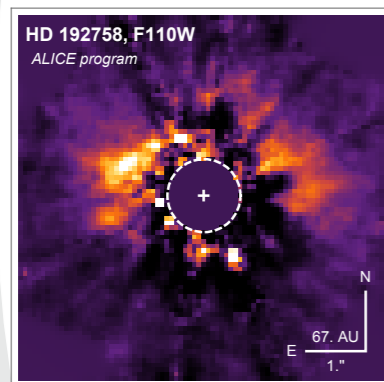
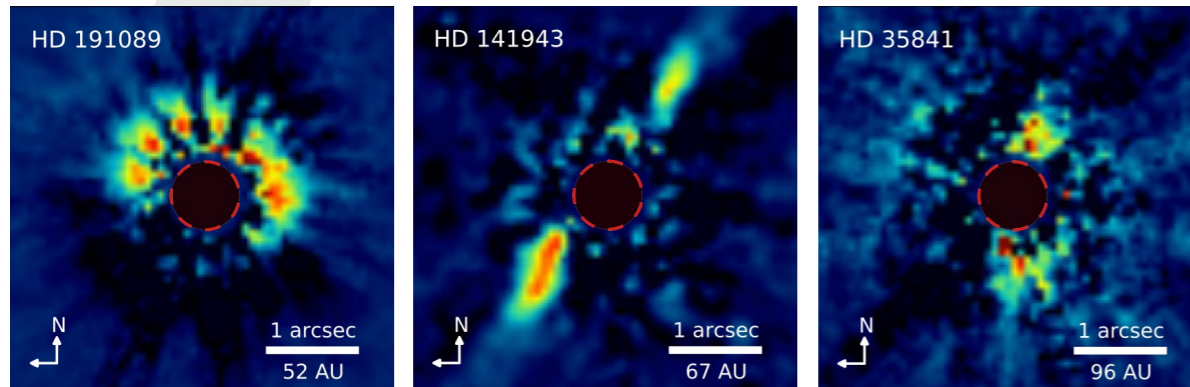


In λ/D
units

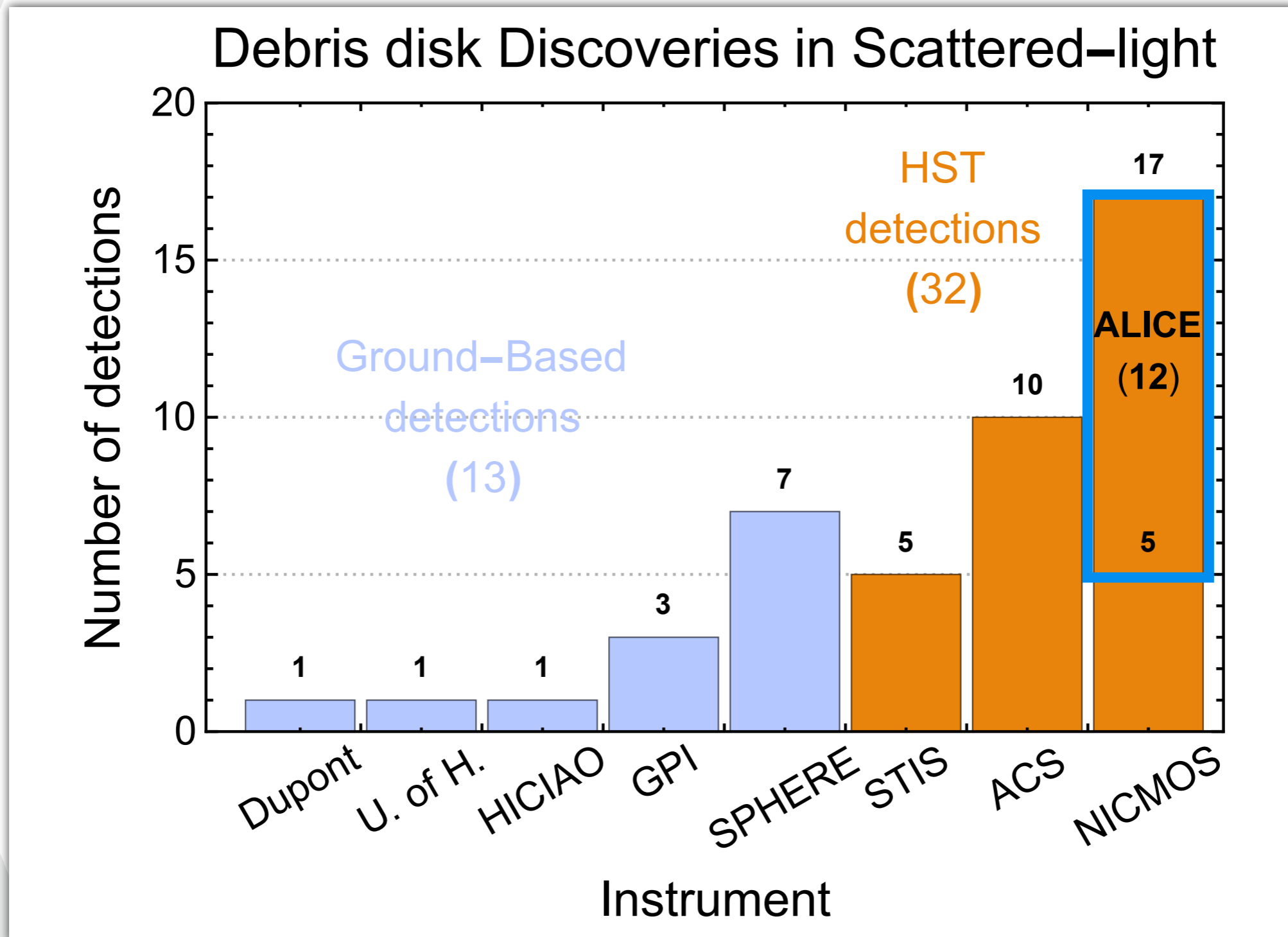
Soummer et al. 2012
Choquet et al. 2014
Hagan et al. 2018

ALICE: Debris disk gallery

Soummer et al. 2014
Choquet et al. 2016
Choquet et al. 2017
Choquet et al. 2018
Marshall et al. 2018



ALICE: Debris disk gallery

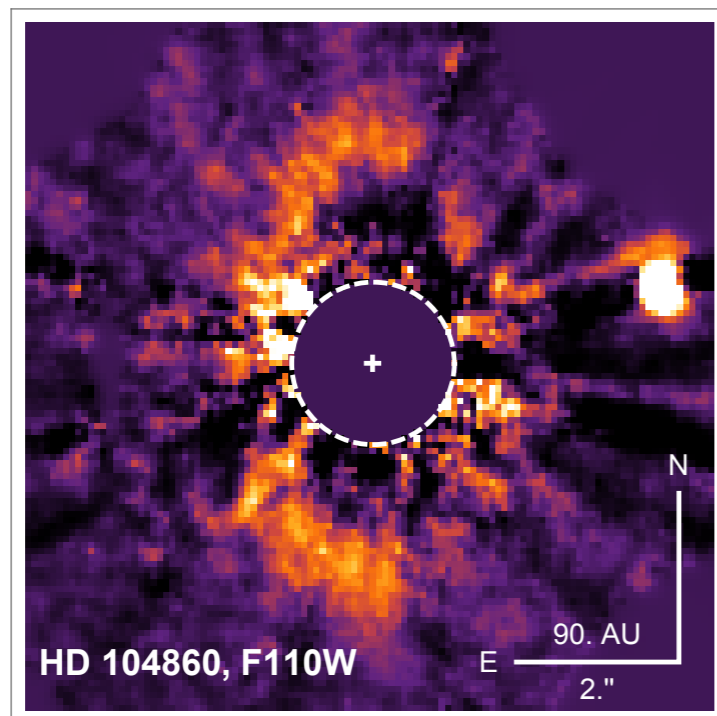


2. Understand the properties of Debris Disks

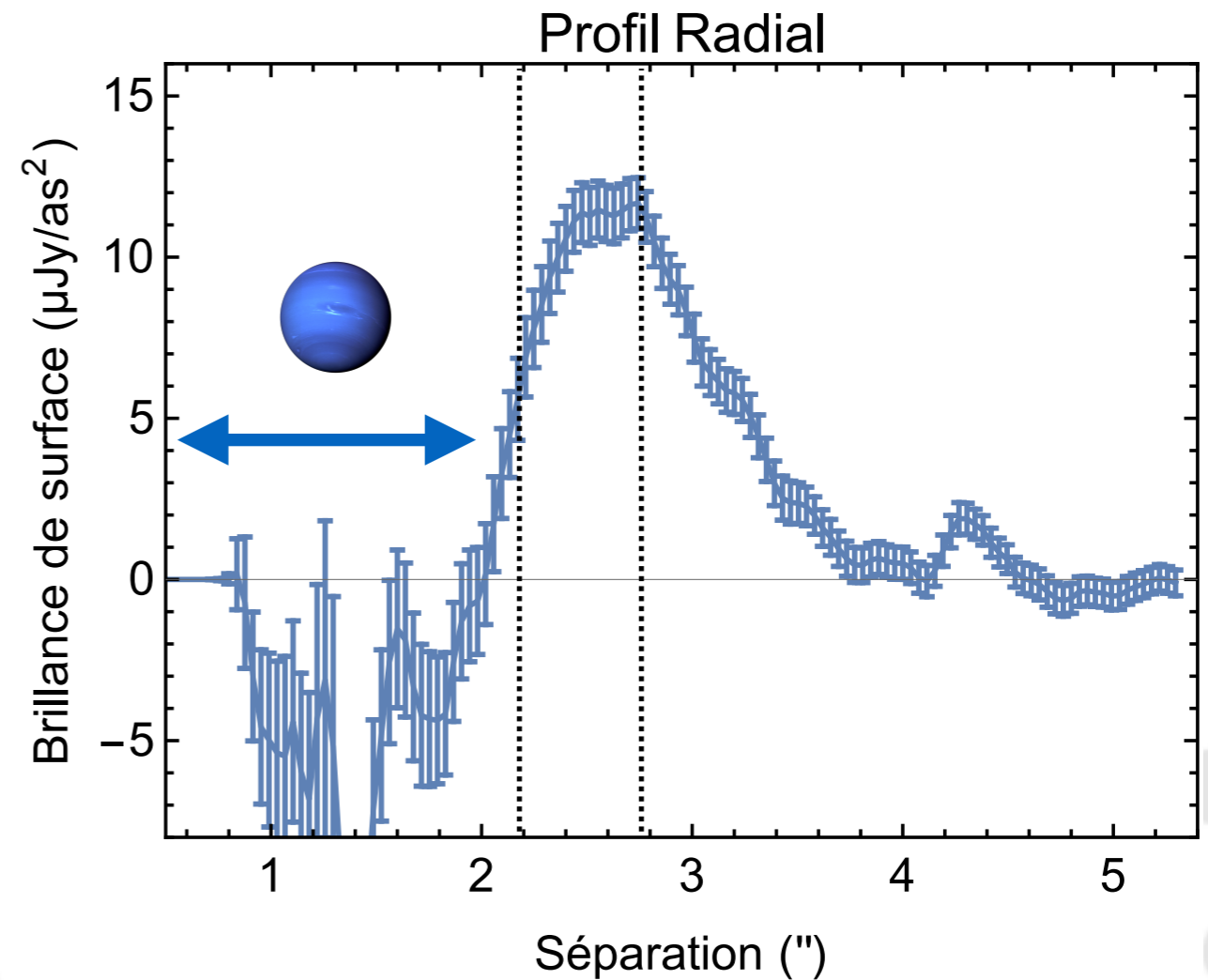
Disk morphologies to probe dynamics

Inner Clearings

Evidence of planets



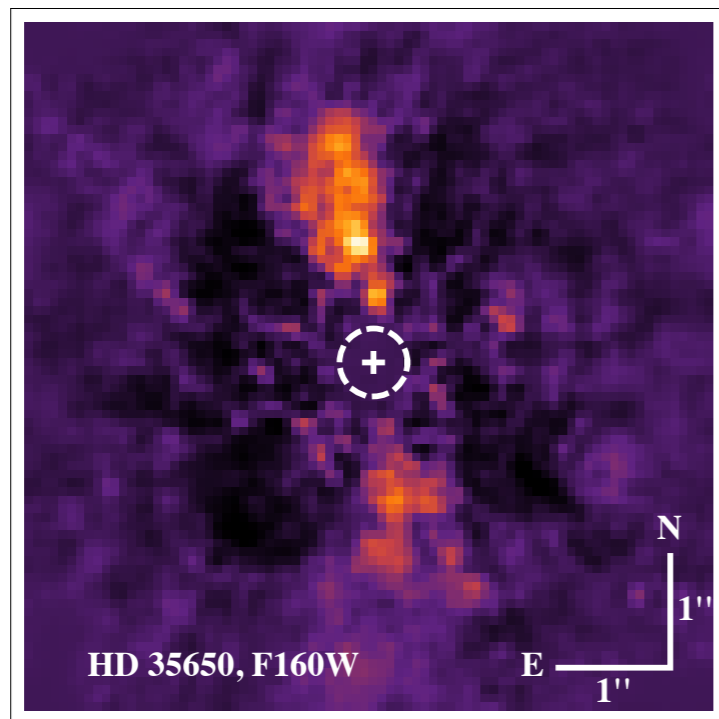
Choquet et al. 2018



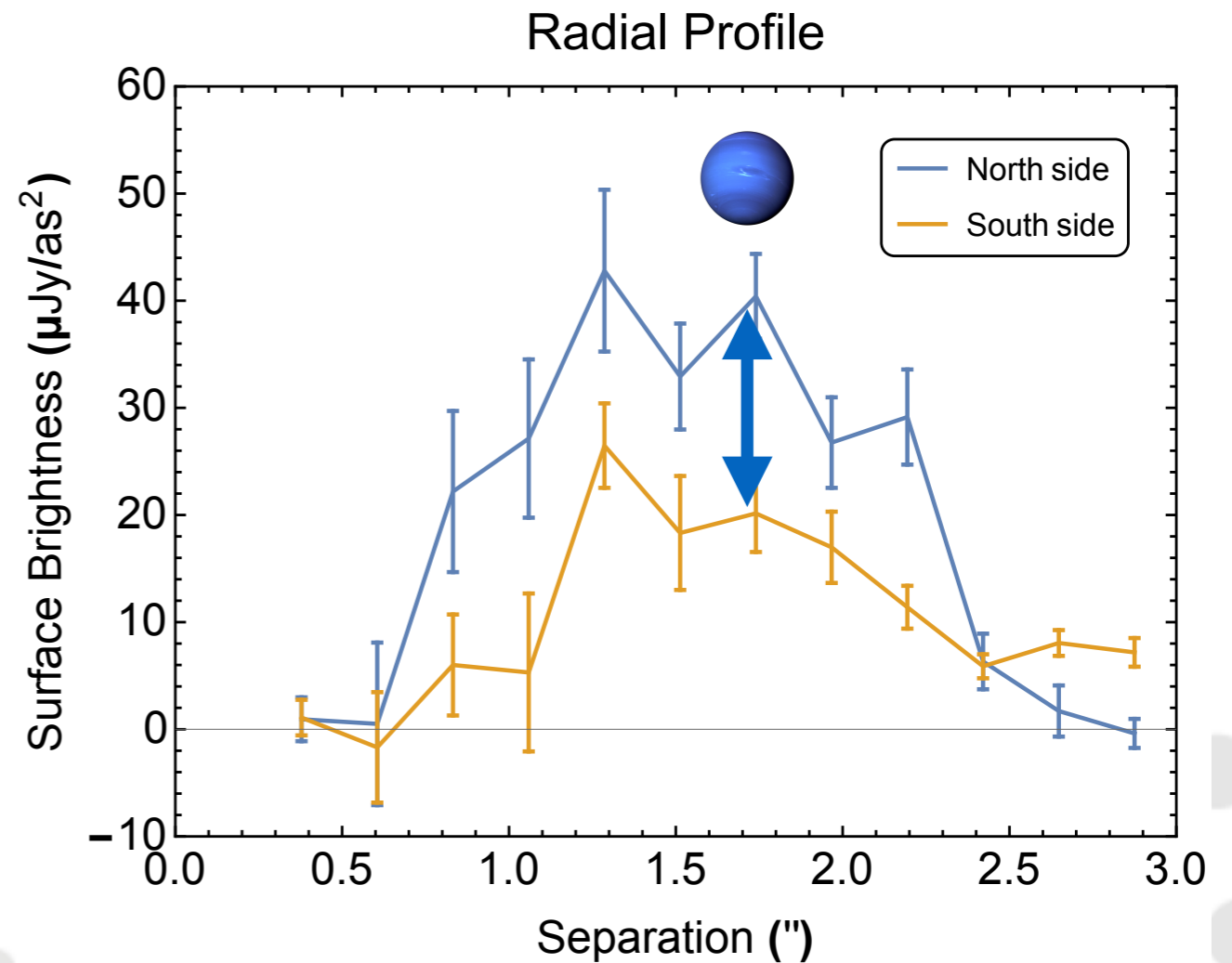
Disk morphologies to probe dynamics

Brightness asymmetries
Offsets from the star

Evidence of planets

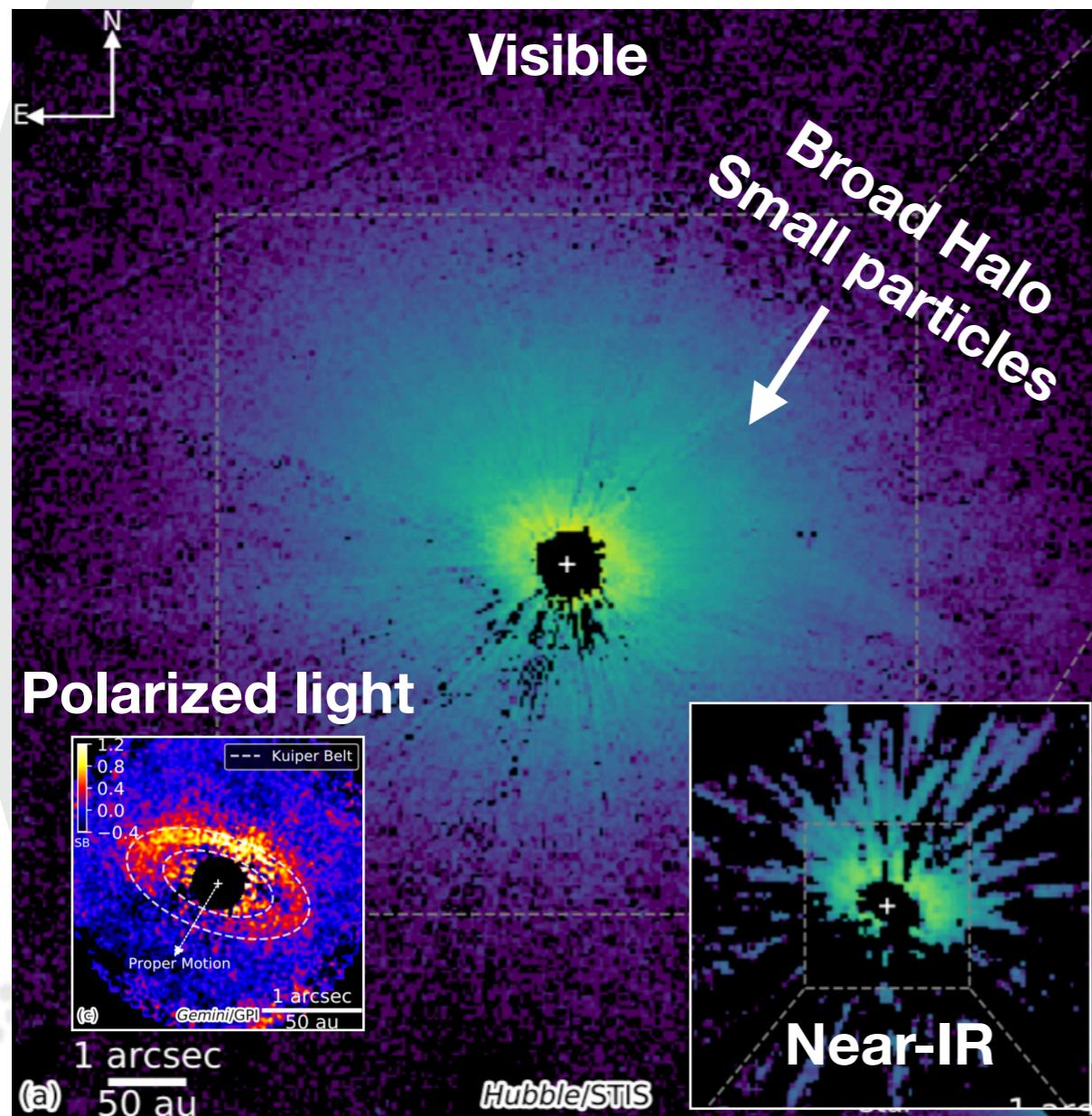


Choquet et al. 2016



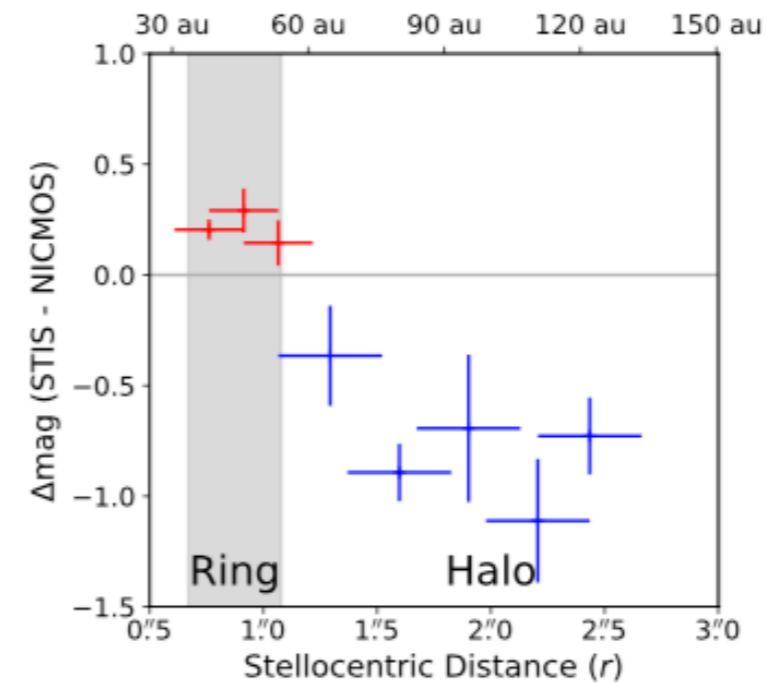
Disks brightness to probe their composition

Color & Scattering phase function

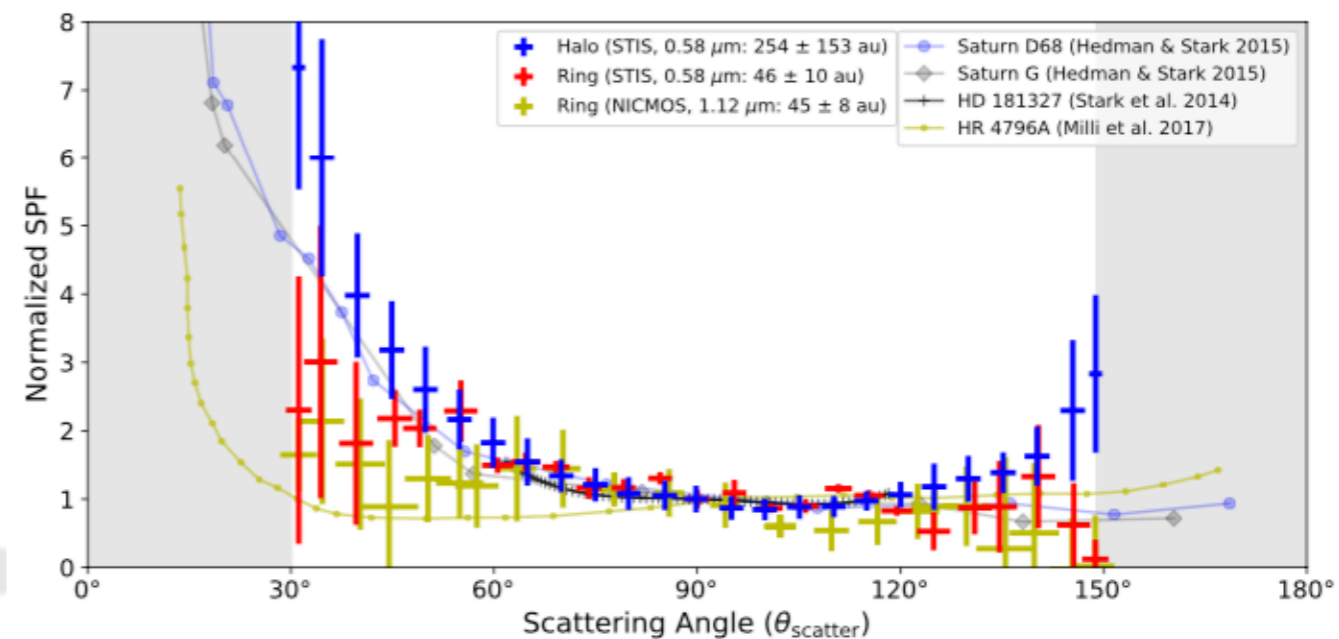


Ren, Choquet et al. 2019, sub.

Vis-nIR Color

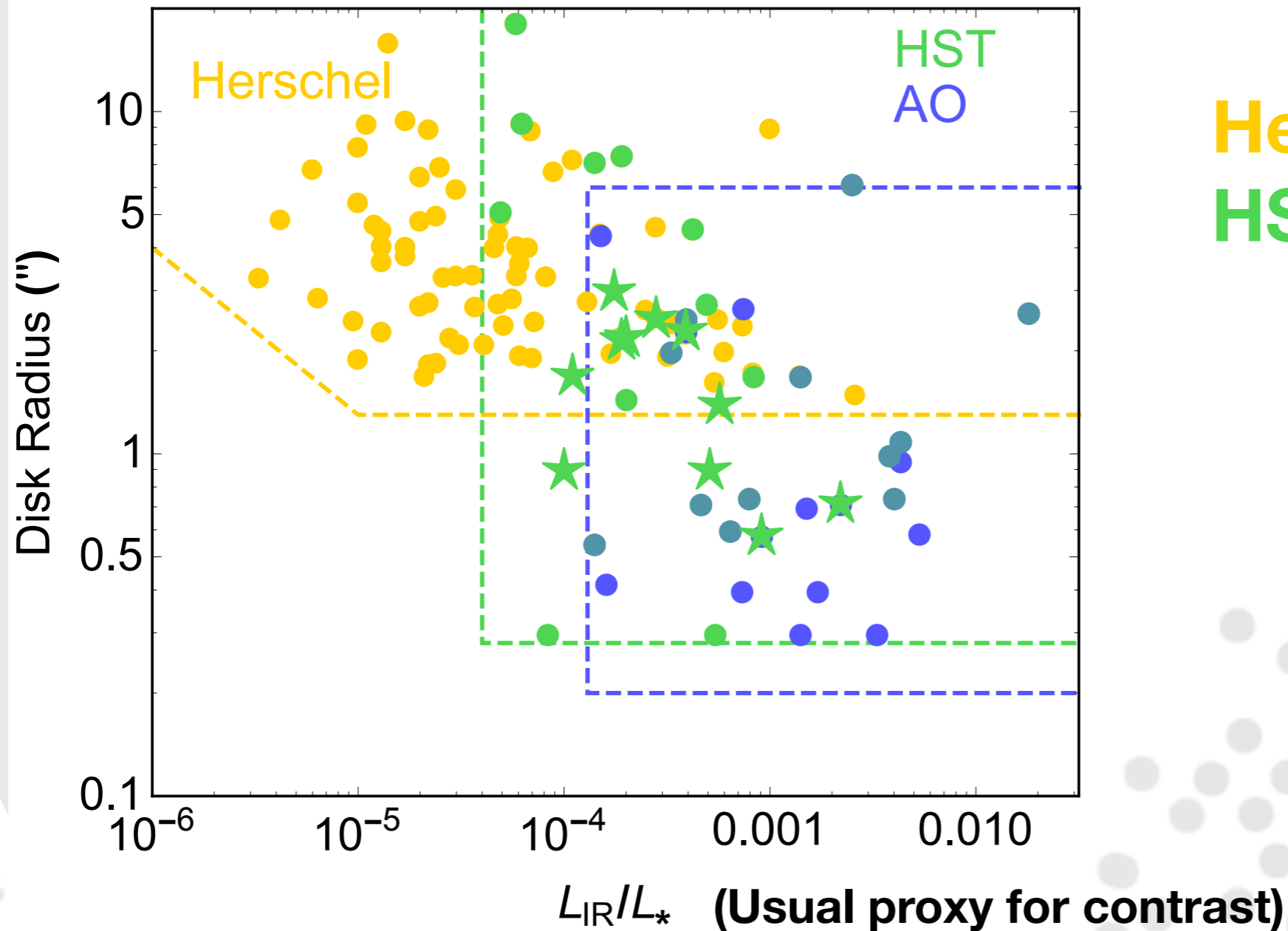


Scattering phase function



Resolved disk population study

Debris disks resolved to date



Herschel: 82

HST & AO: 45

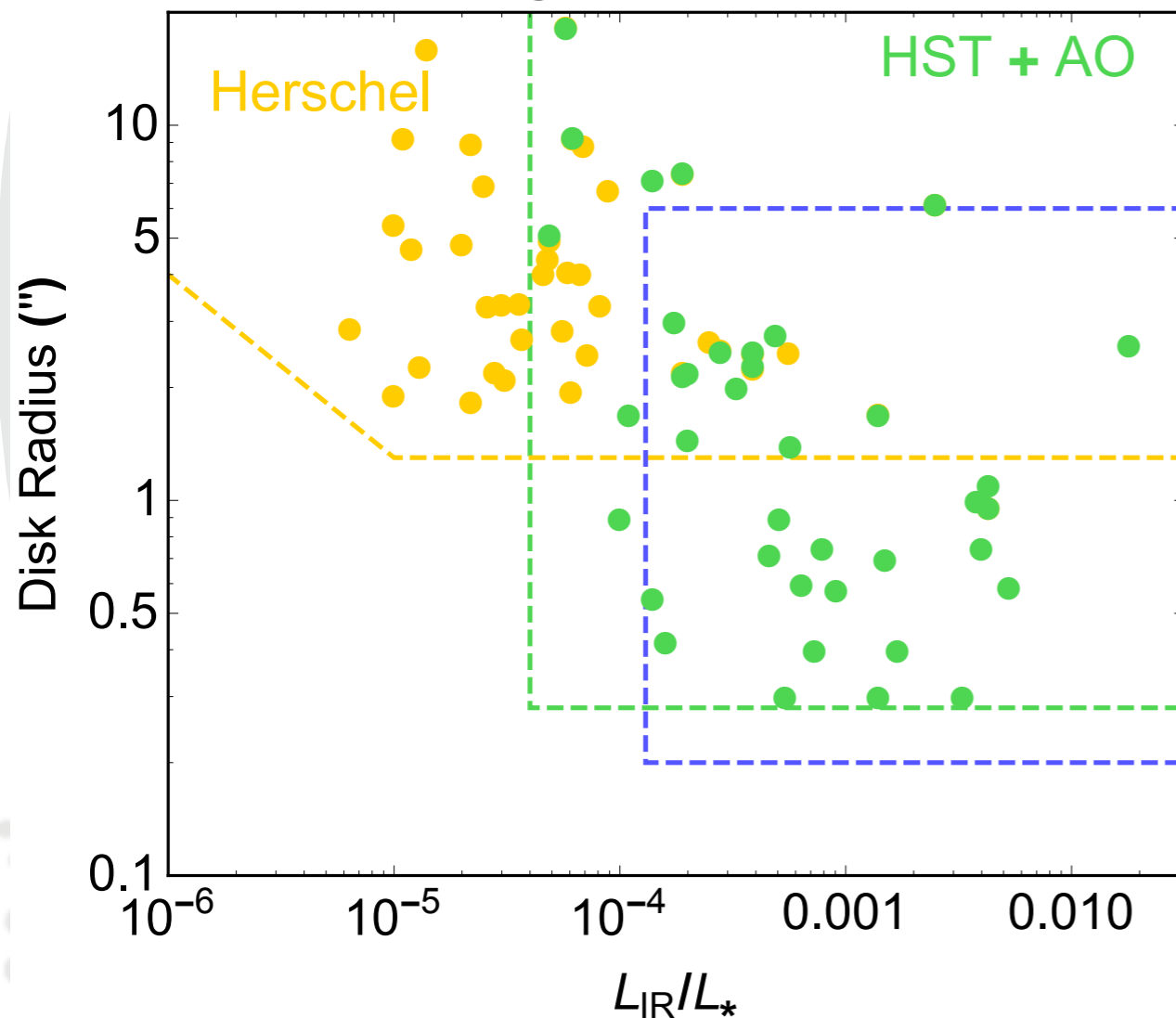
Resolved disk population study

Herschel: 50%
HST + AO: 87%

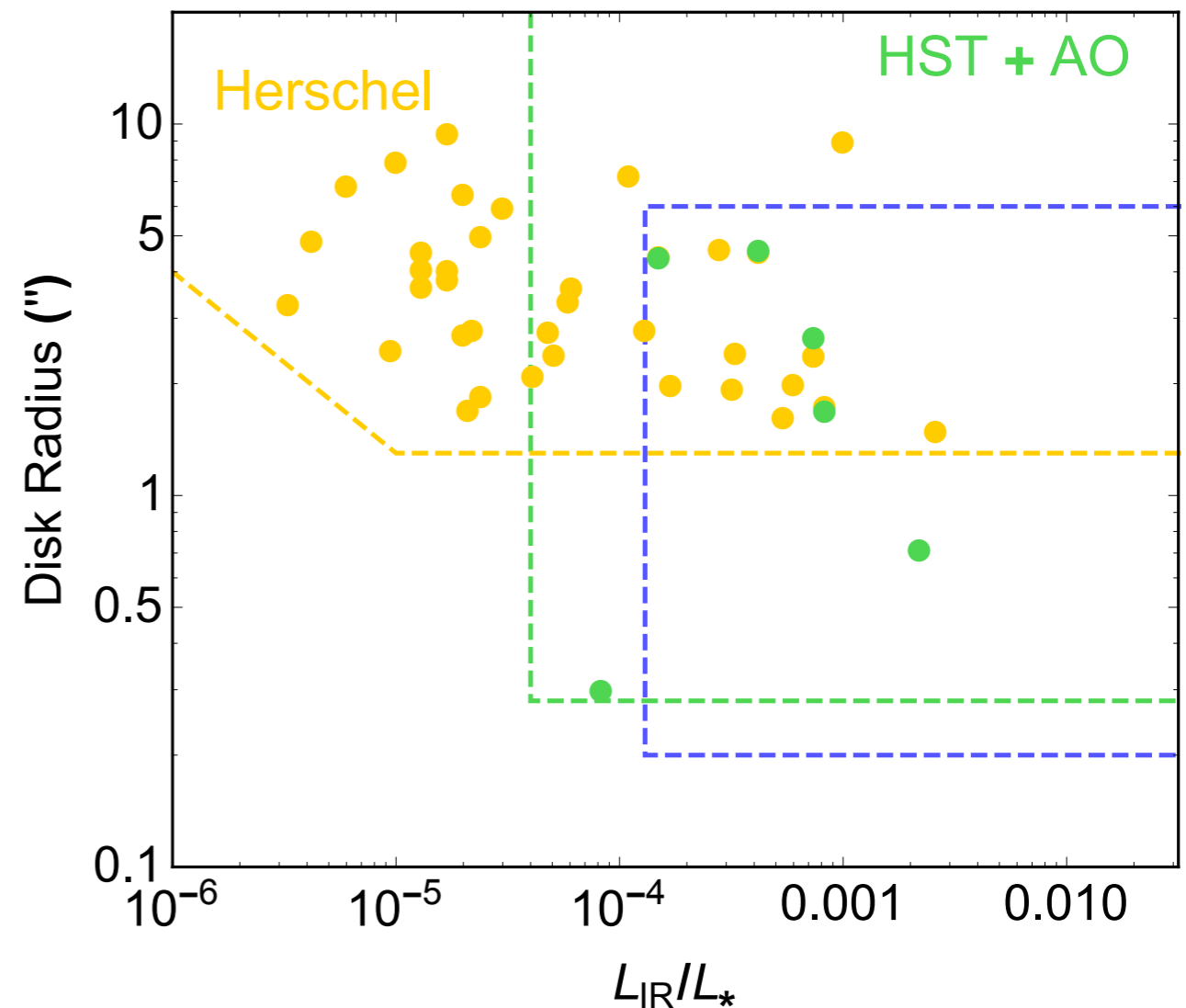
> 45°
inc. >

Herschel: 50%
HST + AO: 13%

High Inclination



Low Inclination

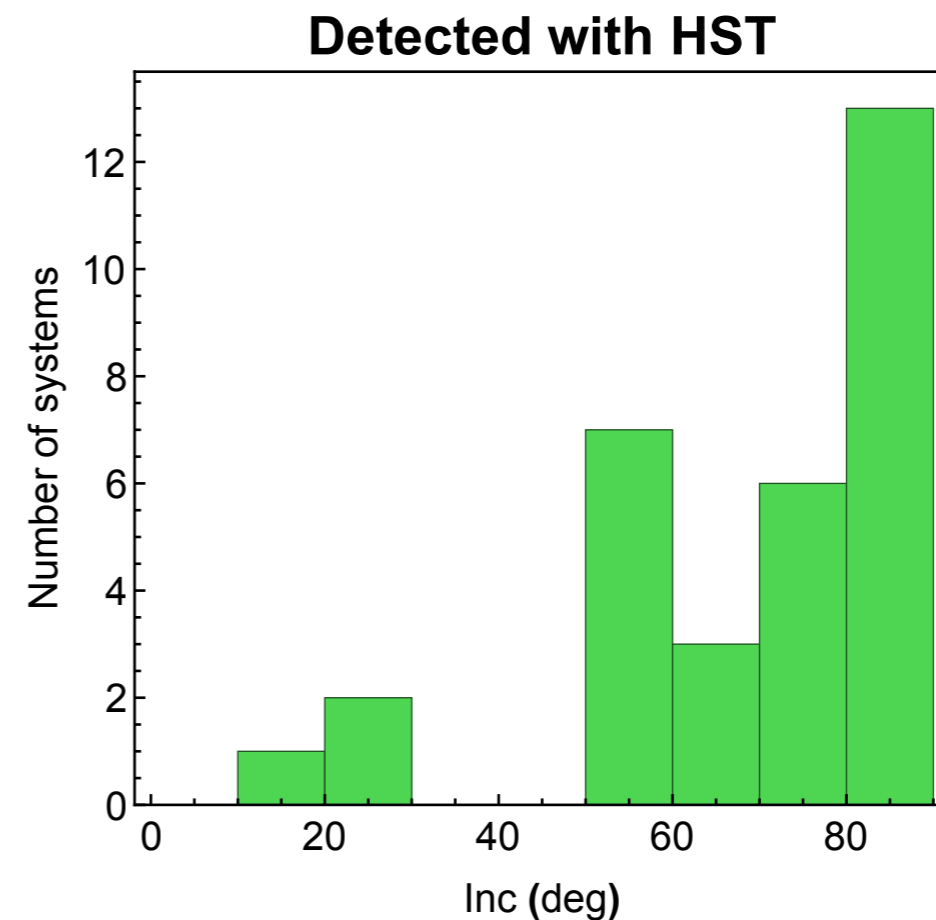
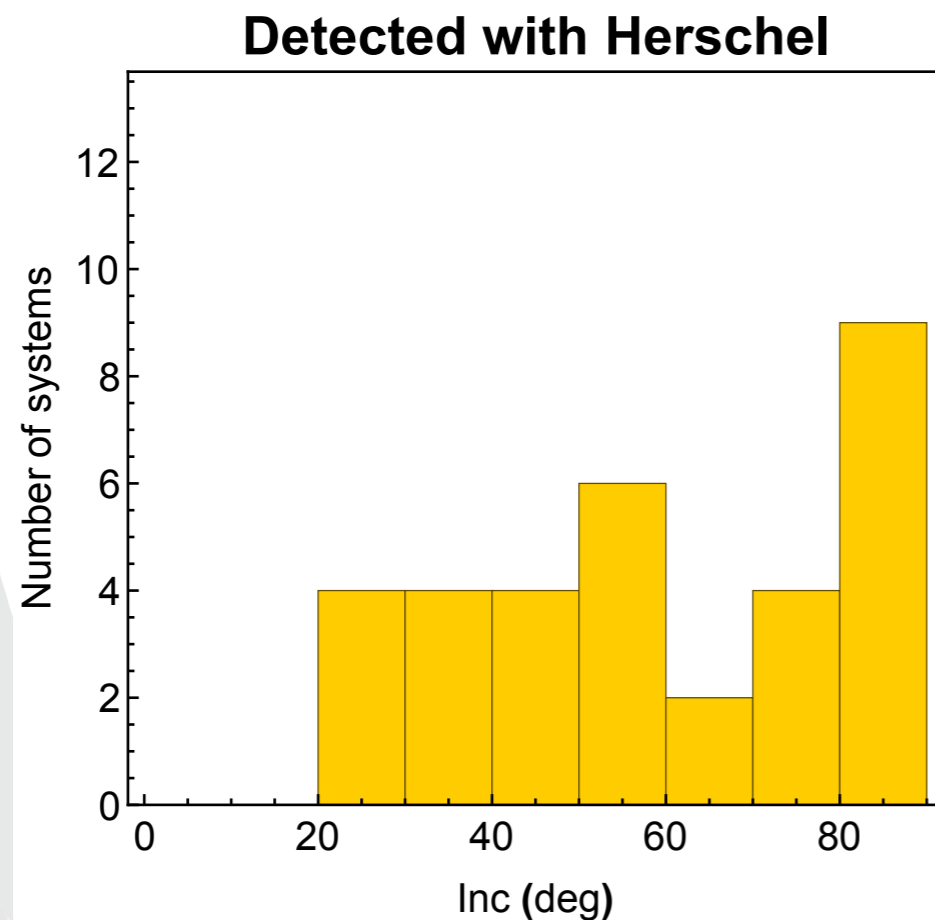


Resolved disk population study

On-going work!

- ▶ MCMC population synthesis
- ▶ Effect of dust scattering properties

Sub-sample: systems observed with HST + RDI



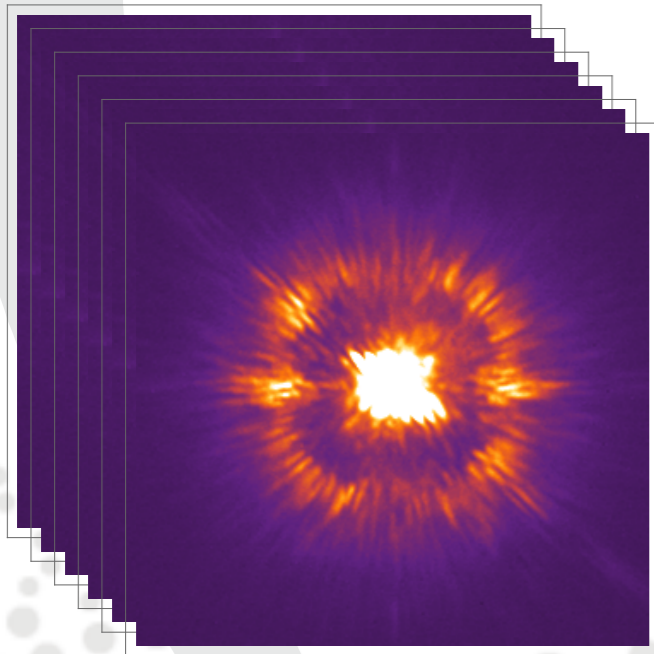
Choquet & Millar-Blanchaer in prep.

Conclusion and Prospects

Image Processing

Reprocess the SPHERE archive

- ▶ Using the AO telemetry
- ▶ Using Machine Learning approaches



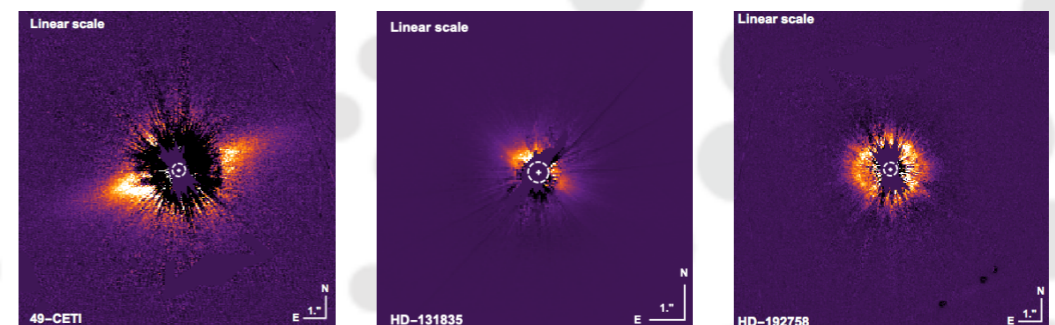
Debris Disk Studies

Characterization of known disks

- ▶ On-going HST program
- ▶ Visible-nIR characterization

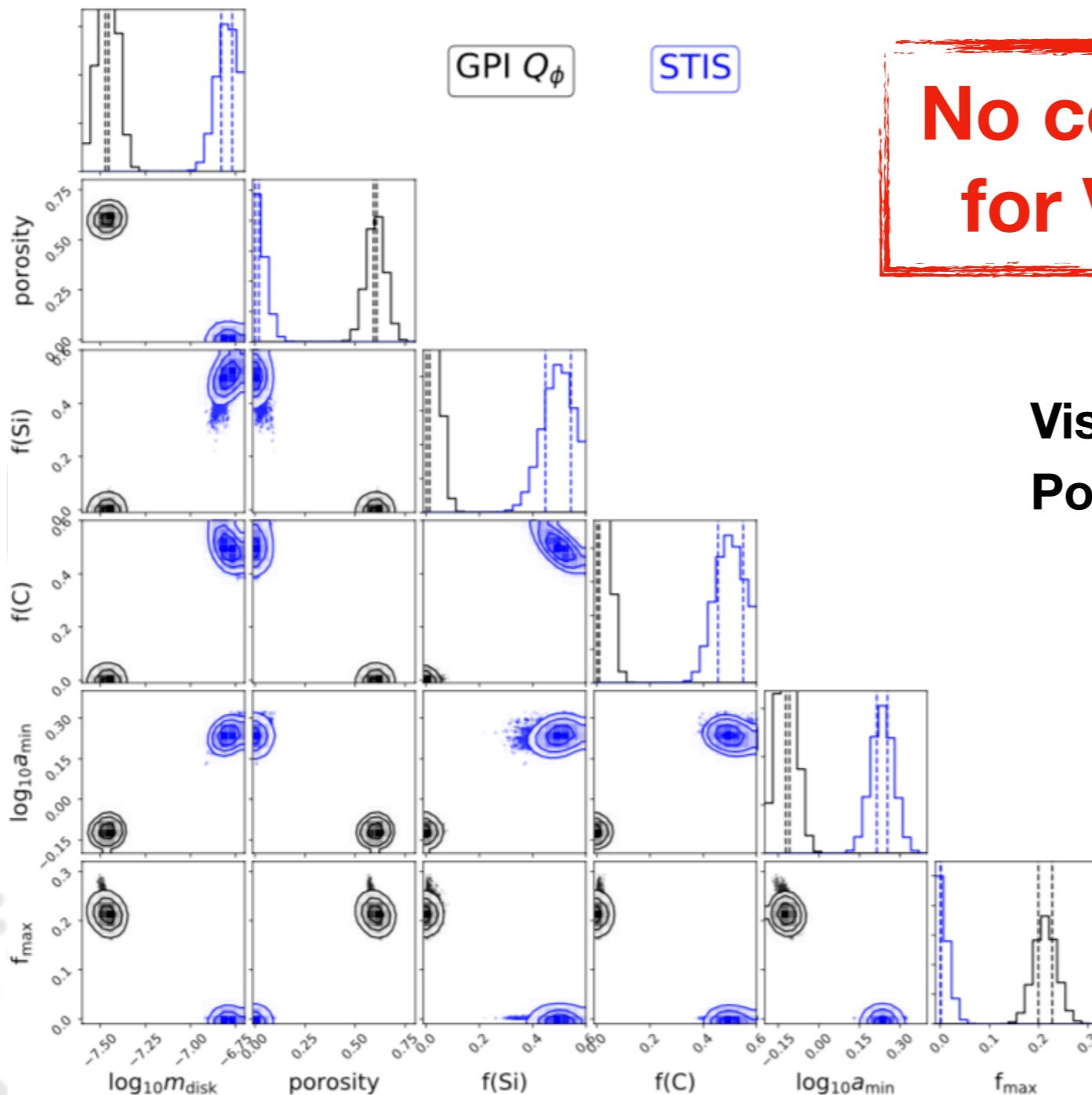
HST Survey of Herschel-resolved disks

- ▶ Improve statistical significance
- ▶ Statistical study of other properties



END

Limitations of the dust models



**No consistent model
for Vis. & Pol. data**

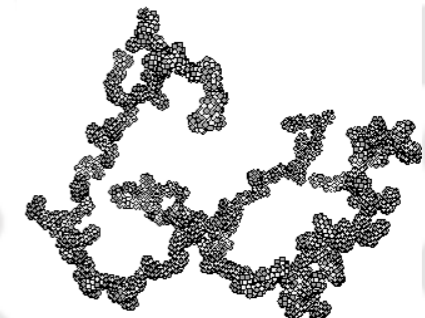
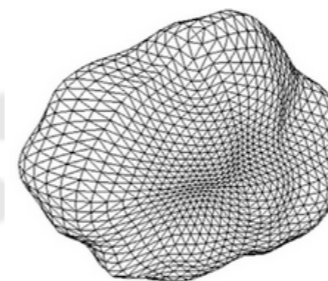


Vis. image: homogeneous sphere

Pol. image: Hollow sphere distribution



Complex shapes, aggregates ?



Ren, Choquet et al. 2019, sub.