A Simulator-based Autoencoder approach for Focal-Plane Wavefront Sensing:

principle, vortex phase diversity, perspectives for on-sky test

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Context	Vortex phase diversity	Simulator-based autoencoder	Extending the method	Application to real data	Conclusions
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Exoplanet imaging

- Limitations:
 - \star small angular separations
 - * high contrasts
- Instrumental solutions:
 - \star adaptive optics
 - \star coronagraphy
- Quasi-static speckles remain.



Martinez et al. 2013



Credit: GMT



Credit: ESO

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Non-common path aberrations



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Focal-plane wavefront sensing



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Limitation: phase sign ambiguity

$$\Theta_{pupil} = \sum_{i=0}^{N} c_i Z_i$$

• Problem: sign ambiguity for even modes.

$$\mathscr{F}\{E_{pupil}(x)\} = \mathscr{F}\{E_{pupil}^*(-x)\}$$

Solution: phase diversity.



in-focus PSF

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out-of-focus PSF

 \rightarrow Decomposition of Θ_{pupil} into Zernike modes



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Vortex coronagraphs



Huby et al. 2015

- Vector Vortex coronagraph (VVC): (Mawet et al. 2005)
- Conjugated phase ramps $e^{i\pm\ell_p heta}$.
- Split circular polarization states and use two in-focus PSFs.



Delacroix et al. 2013



NASA/JPL-Caltech/Palomar Observatory

- Scalar Vortex coronagraph (SVC): (Ruane et al. 2019)
- Same phase ramp *e<sup>iℓ_pθ* for both circular polarization states.
 </sup>
- Use one in-focus PSF.

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Vorte×	phase diversit	.y				
	0	Φ		$\Phi'_{even} = -6$	Concept introd Riaud et al. Φ_{even}	uced by 2012
	in-focus out-of-for $\pm \ell_p$	$\underbrace{\begin{array}{c} \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\$	n-focus	$\underbrace{ \begin{array}{c} \text{out-of-focus} \\ \hline \hline \hline \hline \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ $	Vector VC classical phase diver	rsity
	$+\ell_p$		in-focus	in-focus −ℓ _p	Vector VC polarization separation Or Scalar VC	on

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Deep Convolutional Neural Networks

 \rightarrow Motivation: fast predictions, higher performance, better robustness.

U-Net (Ronneberger et al. 2015):



EfficientNet (Tan et al. 2019):



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Zernike modes reconstruction



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Phase retrieval performance



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Simulator-based Autoencoder (SimAE)



CNN loss function (supervised):

$$\mathcal{L}_{CNN}(z,\widehat{z}(x;\phi)) = \sqrt{\frac{1}{N}\sum_{i}^{N}(z_{i}-\widehat{z}_{i}(x;\phi))^{2})}$$

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Simulator-based Autoencoder (SimAE)



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Simulator-based Autoencoder (SimAE)



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SimAE loss function (unsupervised) \rightarrow Poisson distributions:

$$\mathcal{L}_{\textit{SimAE}}(x;\phi) = -\mathbb{E}_{x \sim \rho(x)} \left[\log \left(\frac{\lambda(x;\phi)^x}{x!} \exp(-\lambda(x;\phi)) \right) \right]$$

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Simulator-based Autoencoder (SimAE)



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SimAE: Performance

- WFE \in [0, 350] nm rms
- SNR $\simeq 100$
- 10^5 training & 10^2 test samples

Evaluated on specific WFE distributions:



Entire WFE distribution:



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SimAE: fit on-the-fly

- Pre-trained on 10⁵ samples
- Fine-tuned on 1 test sample
- Convergence time: \sim 10 seconds



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Improving the simulator					

- Learnable instrumental parameters: for now Vortex rejection factor
- Including AO telemetry into the simulator:



• Use optical propagation package: HEEPS (PyTorch) or δ Lux (Jax)

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Variati	onal inference				

• Add Posterior q(z|x) and Prior distributions p(z):



- KL divergence loss term: $-\beta KL(q(z|x;\phi)||p(z)) \rightarrow \text{training now stable}$
- Evaluation metric: $\mathbb{E}_{p(x)}[\log q(z|x)]$
- How do we pick the best z?



Credits: Jyotirmay Paul

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Subaru/SCExAO instrument



Credits: Subaru/SCExAO and Jyotirmay Paul

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Transfer learning with SCExAO data





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Conclusions

